GENETIC CLASSIFICATION AND NOMENCLATURE OF LUNAR HIGHLAND ROCKS BASED ON THE TEXTURE AND GEOLOGIC SETTING OF TERRESTRIAL IMPACT BRECCIAS.

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The petrogenesis and the chronology of the primordial crustal rocks of the Moon which potentially jive important evi-dence of planet-forming processes, are still a matter of dispute mainly because to-day these rocks form displaced fragments within complex polygenetic breccia formations of the highlands. A geologically and stratigraphically meaningful interpretation of the chemical, physical and mineralogical properties of lunar highland breccias is largely dependent on a genetically oriented, common classification in which macroscopic and microscopic textural properties are used as prime criteria for classifying each sample. In this respect previous classifications (1-7) are often non-uniform and incompatible with each other so that much confusion among those lunar scientists who are not primarily concerned with the sample petrography, was created in the past. A major obstacle for the development of one common, petrogenetically acceptable rock classification was that different kinds of criteria for classification were used by different authors or even within one single classification model. This is best revealed by the present nomenclature of breccias in which non-equivalent terms such as genetic ("thermally metamorphosed", "partially molten"), textural ("poiki-litic", "cataclastic"), color ("black and white", "light gray") or petrographic ("micronoritic", "feldspathic") terms were applied within one and the same classification. A great deal of the difficulty to achieve a satisfactory common classification of lunar breccia samples is due to the lack of or the insufficient use of information about the geological setting of a sample, and to the inadequate consideration or disregard of the well-established classification and petrogenesis of terrestrial impact breccias and their spatial relation to the pertinent crater in which they were formed.

The present paper uses data from two different research projects in which the texture and petrography of all major types of lunar breccias (8,9) and of terrestrial impact breccias at the Ries, Rochechouart, Mien, Dellen, Lappajärvi and Sääksjärvi craters located in central Europa and Scandinavia (10-12)were studied by quantitative stereometric analysis. Comparison of both sets of data clearly shows that there is nothing to be observed in the texture of lunar highland rocks which could not be reconciled or would be in conflict with observations made at terrestrial impact breccias. The results of this comparative study combined with the geological model of breccia formation and distribution at terrestrial craters provides a solid basis for a revised classification of lunar highland rocks. Similar to previous classifications (e.g. 1,2) it is built on the unifying genetic concept that the observed rock textures are predominantly produced by impact-induced processes such as

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melting, fracturing, mechanical mixing and transportation and that the primary rock material (target material) has been affected by such processes repeatedly.

In our proposal (Table 1) each class of breccia established on the basis of textural parameters gets a genetic connotation with respect to the analoguous terrestrial breccia class. This genetic system is irrespective of the multiplicity of brecciation events observed in a particular sample. It concerns only the last impact event which forms the "youngest" matrix of the breccia in question. The petrogenetic history of the various breccia classes is illustrated in Fig.1 and 2. It is proposed to use the terms "impact cataclasite", "impact melt rock", "suevitic impact breccia", "clastic (or fragmental) impact breccia" and "regolith breccia" to denominate the five broad classes of lunar breccias. Within these classes further subdivisions may be made on the basis of the textural characteristics of the matrix (9). Also the terms "crystalline matrix breccia" (2) and "clastic matrix breccia" continue to be useful for the denomination of a subgroup of the impact melt rocks and of the combined two classes of breccias with clastic matrix, respectively. The crystalline matrix breccias may be subdivided most easily by using adjectives such as equigranular, subophitic, poikilitic, or granoblastic in connection with a grain size measure to designate the matrix type (9, 1, 2).

The misconception of large-scale thermal metamorphism (3, 2, 1) which was believed by these authors to affect formations of clastic or glassy matrix breccias to form crystalline matrix breccias is excluded explicitely from this genetic classification. Observations at terrestrial craters indicate that thermally induced solid state recrystallization is restricted to lithic clasts incorporated into hot impact melts. Also the breccias with granulitic matrix (13) could be formed in this way, Granulitic textures alone - so far observed in rather small samples - do not necessarily require a large scale granulite metamorphism as postulated by (13).



Fig. 1

Fig. 2

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<pre>1 370, Part 2, Washington, D.C. ion 370, Part 1, Washington 623, (4) Wilshire, H.G. and imonds, C.H. (1974), Proc. coc.Lunar Sci.Conf. 4th, 481, cd, 753, (8) Stöffler, D.et al. ce, Houston, Texas, (9) cience X (this volume), (10) J. Roddy et al., eds.), 343, Savarica 75, 163, Munich, (this volume), (13) Warner, J.L</pre>	elation to the geological	Relation to the crater topogr.	and the becoment and adouted	crater pasement and ejected . Tegablocks		continuous crater deposits	inside and outside the crater	continuous or discont. upper- most immact formation inside	and outside the crater	ejecta of the uppermost im-	pace to crater (ballistic	transport) large, sheet-like bodies	mostly in central part of cra-	ballistic transport) or intru-	sions into basement rocks	inside and outside of small	sized craters located within the lunar regolith	(with secondary mass transport class.
<pre>James, 0.B., (1977) NASA Special Publication W.C. et al. (1977), NASA Special Publication Scher, J.L. (1972), Proc.Lunar Sci.Conf. 3rd 372), U.S.Geol.Survey Prof.Paper 785, (5) Si 5th, 337, (6) Warner, J.L. et al. (1973), P W.v. et al. (1972), Proc. Lunar Sci.Conf. 3 Hence IX, 1116, Lunar and Planetary Institu Stöffler, D. (1979), Lunar and Planetary Si (1977) in Impact and Explosion Cratering (D (11) Stöffler, D. et al. (1977), Geologica cal. (1979), Lunar and Planetary Si co.Lunar Sci.Conf. 8th, 2051</pre>	land rocks and their re	Terrestr. equivalents		cataciastic target works	10010	"Bunte breccia" at (1)	Ries crater	(2) (2)	V-V ANA V-V	quenched impact melt	particies (e.g.g.giass bombs in suevite)	transifions	· · · · · · · · · · · · · · · · · · ·	conerent crystallized impact melt rocks or	dikes of melt	попе		e mode of emplacement (oulders belong to this
	ication of lunar high	Lunar rock examples	8451	cataclastic ANT-		"light matrix" (3)	breccia (in part)	Apollo 14	WILLE LOCKS	fragment-laden	glasses	Fra Mauro breccias	light-, blue-, green-	gray br., poikil.		regolith breccias		I subseq. ground surg) many of the large b
References. (1) 637, (2) Phinney 0.C., 91, (3) Wal Jackson, E.D. (1 Lunar Sci.Conf. (7) Engelhardt, ((1978), Lunar Sci Knöll, HD. and Pohl, J. et al. (Pergamon, N.Y., et al. (1977), Pl et al. (1977), Pl	Table 1: Classifi 2014:00	Breccia type	monomict	cataclastic	<u>rgueous rocs</u> nolvmict and allo	breccia with	clastic matrix	breccia with	cl. matrix and melt particles	breccia with	glassy matrix	transıtıons	crystall.matrix	1		breccia with cl.	matrix bound by	(2) ballistic and for (1)), (3)

CLASSIFICATION OF LUNAR HIGHLAND ROCKS

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